

# **A Consistent Treatment of Microwave Emissivity and Radar Backscatter for Retrieval of Precipitation over Water Surfaces**

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## ABSTRACT

15 The Global Precipitation Measurement satellite's Microwave Imager (GMI)  
16 and Dual-frequency Precipitation Radar (DPR) are designed to provide the  
17 most accurate instantaneous precipitation estimates currently available from  
18 space. The GPM Combined Algorithm (CORRA) plays a key role in this pro-  
19 cess by retrieving precipitation profiles that are consistent with GMI and DPR  
20 measurements; therefore it is desirable that the forward models in CORRA  
21 use the same geophysical input parameters. This study explores the feasi-  
22 bility of using internally consistent emissivity and surface backscatter cross  
23 section ( $\sigma_0$ ) models for water surfaces in CORRA. An empirical model for  
24 DPR Ku and Ka  $\sigma_0$  as a function of 10m wind speed and incidence angle is  
25 derived from GMI-only wind retrievals under clear conditions. This allows  
26 for the  $\sigma_0$  measurements, which are also influenced by path-integrated atten-  
27 uation (PIA) from precipitation, to be used as input to CORRA and for wind  
28 speed to be retrieved as output. Comparisons to buoy data give a wind rmse  
29 of 3.7 m/s for Ku+GMI and 3.2 m/s for Ku+Ka+GMI retrievals under precip-  
30 itation (compared to 1.3 m/s for clear-sky GMI-only), and there is a reduction  
31 in bias from the GANAL background data (-10%) to the Ku+GMI (-3%) and  
32 Ku+Ka+GMI (-5%) retrievals. Ku+GMI retrievals of precipitation increase  
33 slightly in light ( $< 1$  mm/hr) and decrease in moderate to heavy precipitation  
34 ( $> 1$ mm/hr). The Ku+Ka+GMI retrievals, being additionally constrained by  
35 the Ka reflectivity, increase only slightly in moderate and heavy precipitation  
36 at low wind speeds ( $< 5$  m/s) relative to retrievals using the surface reference  
37 estimate of PIA as input.

## 38 1. Introduction

39 Algorithms for estimating precipitation from space-borne radars at attenuating frequencies (e.g.,  
40 TRMM PR (Iguchi et al. 2000, 2009), CloudSat (Mitrescu et al. 2010), GPM DPR (Grecu et al.  
41 2011)) have long realized the benefit of an estimate of the path-integrated attenuation (PIA) that is  
42 independent of the reflectivity profile for the purposes of constraining the integrated and surface  
43 precipitation amount. In general, such an estimate of the PIA is obtained via a form of the sur-  
44 face reference technique (SRT; (Meneghini et al. 2000, 2004)), which subtracts the surface radar  
45 backscatter cross-section ( $\sigma_0$ ) in a precipitating column from a precipitation-free reference. The  
46 difference is then assumed to be due to attenuation from precipitation after accounting for multiple  
47 scattering (Battaglia and Simmer 2008) and the effect of precipitation on the surface itself (Seto  
48 and Iguchi 2007). If the ratio of this difference to the uncertainty in the reference value, known  
49 as the reliability factor, is large, then the precipitation retrieval is more strongly constrained, be-  
50 cause the PIA is sensitive to the vertically-integrated third moment of the particle size distribution  
51 whereas the reflectivity is sensitive to the sixth moment.

52 Algorithms that make simultaneous use of passive microwave and radar data (Haddad et al.  
53 1997; Grecu et al. 2004; Munchak and Kummerow 2011) generally use the SRT PIA along with  
54 microwave radiances to constrain the precipitation profile (indeed, PIA can be the dominant con-  
55 straint because of its high resolution relative to the passive microwave footprint, especially when  
56 the reliability factor is large). These algorithms also require knowledge of the surface emissiv-  
57 ity in order to forward model the brightness temperatures for comparison to observations. Since  
58 emission and reflection are related processes, it is logical for a combined algorithm to exploit  
59 any relationships between  $\sigma_0$  and emissivity that may exist. Over water surfaces, it is known  
60 that wind-induced surface roughness and foam have a large impact on  $\sigma_0$  and emissivity; thus, it

61 should benefit a combined algorithm to retrieve the 10m wind speed in order to achieve internal  
62 consistency between the forward-modeled PIA and brightness temperatures.

63 The purpose of this work is not only to highlight the benefits of unifying the active and passive  
64 surface characteristics for the purpose of precipitation retrievals from GPM, but also to demon-  
65 strate the feasibility of combined DPR-GMI retrievals of surface wind over water, particularly  
66 when precipitation is present. This has historically been problematic for both passive and active  
67 (scatterometer) wind retrievals (Weissman et al. 2012), despite the high motivation to develop ca-  
68 pabilities to monitor the strength of tropical and extratropical cyclones. For passive measurements,  
69 higher frequency channels ( $> 19$  GHz) can become opaque to the surface in rain and clouds, and  
70 although the surface emission is not fully obscured at lower frequencies, measurements at multiple  
71 frequencies near the C-band are required to distinguish the surface and rain column contributions  
72 to the observed radiances (Uhlhorn et al. 2007). However, the large footprints that are character-  
73 istic of spaceborne microwave radiometers at these frequencies are not optimal for retrievals of  
74 wind and precipitation due to non-uniformity within the footprint. Even outside of rain, cross-  
75 talk between wind, water vapor and cloud liquid water can bias wind retrievals (O'Dell et al.  
76 (2008); Rapp et al. (2009)). Also, rain creates an additional source of surface waves, which can  
77 either enhance or damp surface backscatter, depending on angle, frequency, and wind speed (Stiles  
78 and Yueh (2002), Seto and Iguchi (2007)). Backscattering from the rain itself can also enhance  
79 the measured surface cross-section, particularly for scatterometers that are designed to maximize  
80 signal-to-noise ratio by employing relatively long pulse widths and large footprint sizes (Li et al.  
81 2002). Finally, in high winds the sensitivity of  $\sigma_0$  to wind speed is low (Donnelly et al. (1999);  
82 Fernandez et al. (2006)), limiting the accuracy of retrievals even if rain effects are accounted for.

83 As of yet, only the short-lived Midori-II AMSR-SeaWinds combination of passive and active  
84 instruments have been designed specifically for the measurement of ocean winds, but several in-

85 vestigators have taken advantage of existing platforms with these measurements (e.g., TRMM and  
86 Aquarius) or coincident overpasses of scatterometer and passive microwave radiometers to eluci-  
87 date further information about the atmosphere and sea state than is possible from either instrument  
88 type alone. Studies based on the TRMM microwave imager (TMI) and precipitation radar (PR)  
89 have often used the TMI-based wind retrievals as a reference to develop geophysical model func-  
90 tions (GMFs) for PR, which relate wind speed and  $\sigma_0$  (e.g., Li et al. (2002); Freilich and Vanhoff  
91 (2003); Tran et al. (2007)). These are then used to retrieve the wind field independently with PR  
92 (Li et al. 2004) either as a standalone product or for use as a reference to estimate the rain-induced  
93 attenuation as an input to the rainfall estimation algorithms. In the case of WindSat, a comparison  
94 of its retrievals and QuickScat wind vectors in coincident overpasses was performed by Quilfen  
95 et al. (2007), who found differences between the two depended on wind speed and water vapor (a  
96 consequence of the aforementioned cross-talk between parameters). The authors also attempted to  
97 combine the two sets of measurements via multiple regression. They found that adding QuickScat  
98 to WindSat did not improve wind retrievals outside of rain, but they did note a slight improve-  
99 ment under raining conditions. More recently, the Aquarius satellite, which offers active and  
100 passive measurements at L-band for the purpose of ocean salinity retrieval, was launched. Yueh  
101 et al. (2013) developed GMFs based on SSM/I and NCEP reanalysis collocations and found that  
102 the resulting combined active-passive retrievals of wind speed and salinity compared favorably to  
103 salinity retrievals where ancillary data was used to set the wind vector.

104 The growing number of satellites with active and passive microwave instruments (e.g., TRMM,  
105 GPM, Aquarius, SMAP), along with airborne platforms (e.g., the NASA Global Hawk Hurricane  
106 and Severe Storm Sentinel-HS3) represents an opportunity to use these combinations to retrieve  
107 ocean winds, particularly under conditions (such as rain) where single-sensor methods are under-  
108 constrained. This study is based on data from the Global Precipitation Measurement (GPM) satel-

lite, which has a particularly useful set of measurements for developing the GMFs due to the well-calibrated, high resolution GPM Microwave Imager (GMI) instrument (Draper et al. 2015) and a dual-frequency precipitation radar (DPR) which improves the capability to separate surface effects from rain-induced attenuation. Our strategy (Figure 1) is to develop a GMF for DPR based upon co-located GMI wind retrievals, and then use this GMF under raining conditions by modifying the combined GPM DPR-GMI precipitation retrieval algorithm CORRA (Olson and Masunaga 2015). In order to have as accurate a wind reference as possible, we evaluate three emissivity models after calculating offsets under clear and calm conditions to achieve consistency with the GMI calibration. Next, we use all available matchups of GMI and DPR under non-precipitating conditions to develop the GMFs. This process is presented in section 2. Next, the use of GMFs in the GPM combined GMI-DPR ensemble filter retrieval framework, including validation of winds in regions of precipitation against buoy measurements, is described in section 3, followed by a summary in section 4.

## 2. Development of Geophysical Model Functions for DPR

Although physical models exist to describe the relationship between wind speed, the wave spectrum, and backscatter (Durden and Vesecky (1985); Majurec et al. (2014)), the desire for GPM applications is to be as internally consistent as possible between the emissivity model and DPR GMF. Therefore, the strategy in this study is to derive empirical GMFs from clear-sky matchups of DPR and GMI-derived 10m wind retrievals, eliminating as much as possible the error from precipitation and cloud cross-talk described in section 1, then apply those GMFs to retrievals under all conditions. The use of empirical GMFs derived in this manner is standard practice in the scatterometer community (Migliaccio and Reppucci 2006).

131 The first step in this process is to generate the clear-sky wind retrievals and then assess their error  
 132 relative to buoy observations. In the absence of precipitation, the microwave radiances measured  
 133 by GMI are primarily sensitive to the surface emission, atmospheric temperature and water vapor  
 134 profile, and cloud liquid water. These parameters can be solved for using optimal estimation, also  
 135 known as variational, retrieval techniques. These have been implemented for microwave sensors  
 136 by Elsaesser and Kummerow (2008) and Boukabara et al. (2011), and a blend of their approaches is  
 137 used to derive the surface and atmospheric properties from GMI by minimizing the cost function:

$$J = (\mathbf{x} - \mathbf{x}_a)^T \mathbf{S}_x^{-1} (\mathbf{x} - \mathbf{x}_a) + (\mathbf{y} - f(\mathbf{x}))^T \mathbf{S}_y^{-1} (\mathbf{y} - f(\mathbf{x})). \quad (1)$$

138 The components of the optimal estimation retrieval are the state vector ( $\mathbf{x}$ ) and covariance ma-  
 139 trix ( $\mathbf{S}_x$ ), the observation vector ( $\mathbf{y}$ ) and covariance matrix ( $\mathbf{S}_y$ ), and forward model  $f(\mathbf{x})$ . For  
 140 water surfaces, the state vector consists of the 10m wind speed, cloud liquid water path, and a  
 141 set of variables representing the values of the leading empirical orthogonal functions (EOFs) of  
 142 the atmospheric temperature and water vapor profile. These EOFs were derived from 10 years  
 143 of MERRA reanalysis (Rienecker et al. (2011); NASA/GMAO (2008)) independently in 1K SST  
 144 bins. The number of leading EOFs is chosen such that at least 99% of the variance in temperature  
 145 and water vapor is explained by the selected EOFs. The EOFs are used to simultaneously adjust  
 146 the initial atmospheric temperature and water vapor profiles in order to match the observed GMI  
 147 radiances. This is a change from the Elsaesser and Kummerow (2008) method, which assumed a  
 148 constant lapse rate and scale height for water vapor. These assumptions are sufficient for matching  
 149 observations near the 22-GHz water vapor absorption line, where radiances are mostly sensitive to  
 150 the total column-integrated amount of water vapor and are less sensitive to its vertical structure and  
 151 emitting temperature. However both the vertical structure of water vapor and temperature matter  
 152 for modeling the additional channels near 183 GHz on GMI, so some method of adjusting the



153 shape of the profile in mid and upper levels is necessary. The EOFs represent the climatological  
154 co-varying structures in temperature and water vapor profiles, and are a robust way to adjust both  
155 without requiring temperature sounding channels (e.g., 50-55 GHz). The *a priori* (and initial) state  
156  $\mathbf{x}_a$  is the MERRA reanalysis interpolated in time and space to the GMI pixel location.

157 Because the atmosphere is represented by EOFs and no covariance between the atmosphere and  
158 wind/cloud is assumed, the state covariance matrix  $\mathbf{S}_x$  is diagonal. The observation vector consists  
159 of the 13-channel GMI radiances from the GMI Level 1C-R (intercalibrated and co-located) prod-  
160 uct (GPM Science Team 2015). The co-location matches the high-frequency (HF) observations  
161 (166V&H,  $183\pm3$ , and  $183\pm7$  GHz), which are observed at  $49.2^\circ$  earth incidence angle, with the  
162 lower-frequency (LF) observations, which are observed at  $52.8^\circ$  earth incidence angle. A diagonal  
163 matrix for  $\mathbf{S}_y$  is also assumed, with values of instrument noise (Hou et al. 2014) plus additional  
164 error determined from buoy matchups (Table 1) to account for forward model error and inexact  
165 footprint matching.

166 The forward model is derived from the Community Radiative Transfer Model (CRTM) Emis-  
167 sion (non-scattering atmosphere) model, modified to include the downwelling path length cor-  
168 rection for roughened water surfaces as described by Meissner and Wentz (2012) and using the  
169 same atmospheric layers that are provided by MERRA products up to 10 hPA. Absorption by at-  
170 mospheric gases is calculated from Rosenkranz (1998) and Tretyakov et al. (2003). Cloud liquid  
171 absorption follows Liebe et al. (1991) and cloud water is assumed to follow an adiabatic profile  
172 (Albrecht et al. 1990). Since the surface emissivity and its relationship to wind speed is of funda-  
173 mental importance to this study, three emissivity models were tested for their ability to produce  
174 unbiased clear-sky radiances when forced with buoy-observed surface winds (within 30 minutes  
175 of a GPM overpass) and MERRA atmospheric profiles: FASTEM4/5 (as implemented in CRTM;

176 Liu et al. (2011)) and the Meissner and Wentz (2012) (hereafter MW) model.<sup>1</sup> Wind direction  
 177 was not considered in this study as only the MW model is capable of representing wind direction-  
 178 induced emissivity changes. Instead we include the wind-direction induced error in total model  
 179 error which is derived from buoy matchups. The source of wind observations in this study is the  
 180 International Comprehensive Ocean-Atmosphere Data Set version 2.5 (ICOADS; Woodruff et al.  
 181 (2011); NCDC/NESDIS/NOAA (2011, updated monthly)) from April 2014-March 2015. Only  
 182 observations from platforms with a known anemometer height ( $h_b$ ) were considered, and all winds  
 183 were adjusted to 10m assuming neutral buoyancy using the relationship (Hsu et al. 1994):

$$w_{10} = w_b(10/h_b)^{0.11}. \quad (2)$$

184 Before the emissivity models can be intercompared, sensor calibration must be considered. Fol-  
 185 lowing Meissner and Wentz (2012), a calm-wind offset ( $\delta_0$ ) was determined for each emissivity  
 186 model and each GMI channel. These offsets were obtained by first selecting a subset of ICOADS  
 187 observations with 10m winds less than  $3.5 \text{ m s}^{-1}$ , where the emissivity-wind relationship is linear.  
 188 To filter out clouds, observations were excluded if the polarization difference at 89 GHz was less  
 189 than an SST-dependent threshold representing a cloud liquid water path of  $0.01 \text{ kg m}^{-2}$  under  
 190 average atmospheric conditions or the spatial standard deviation (within 15 km) of 89 GHz Tb  
 191 was greater than 2 K. The RTM was then forced with the observed SST and wind speed and in-  
 192 terpolated MERRA atmospheric profile. The offsets were then calculated in order to minimize the  
 193 bias between observed and simulated GMI brightness temperatures. No offsets were applied to  
 194 the 183 GHz channels, as these were not sensitive to the surface emissivity in the matchups. The  
 195 offsets and root-mean-square error (after offsets have been applied) are given for each channel  
 196 and emissivity model in Table 1. The biases are different for each model at low frequencies, but  
 197 similar or identical at 166 GHz, indicating low sensitivity of the brightness temperatures to emis-

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<sup>1</sup>Note that the MW model does not include frequencies higher than 90 GHz and FASTEM5 was substituted at these frequencies.

sivity at these channels and therefore low confidence in the offsets, which are likely influenced by the water vapor absorption model and/or absolute calibration of GMI. The root-mean-square-error (rmse) values, which are not sensitive to the choice of emissivity model, represent the error from other components of the forward model (such as wind direction and water vapor absorption) plus instrument noise, and are used as the diagonal components of  $\mathbf{S}_y$ .

Next, each emissivity model was evaluated under the full range of conditions encountered in the GMI buoy overpasses. The retrieval was performed with each emissivity model and the retrieved winds are compared with observations in Figure 2. These results were filtered to remove precipitation by applying a maximum threshold of 1.0 for the normalized cost function. It is apparent from these results that the MERRA analysis is biased high at observed wind speeds below  $3 \text{ m s}^{-1}$  and biased low above this threshold. The retrievals using the different emissivity models behave similarly to each other up to about  $8 \text{ m s}^{-1}$  and remove most of this bias, but diverge due to different foam models (implicit in MW and explicit in FASTEM 4/5). At observed wind speeds greater than  $15 \text{ m s}^{-1}$ , FASTEM4 begins to diverge below the observed wind speed whereas FASTEM5 diverges above more severely. The MW model gives a slight low bias of as much as  $1 \text{ m s}^{-1}$  at  $10\text{-}15 \text{ m s}^{-1}$  but recovers to near zero at higher speeds. The overall root-mean-squared error in clear conditions for the MW model is  $1.3 \text{ m s}^{-1}$  (equivalent to WindSat) and, because of its low bias over the range of observed wind speeds, is chosen to generate the DPR GMFs.

The DPR GMF was generated by averaging the observed  $\sigma_0$  from the DPR Level 2 product (Iguchi and Meneghini 2014), removing the two-way attenuation from gases and cloud liquid water (which are determined from the GMI retrievals), in wind speed bins with  $0.5 \text{ m s}^{-1}$  spacing from  $0$  to  $10 \text{ m s}^{-1}$ ,  $1 \text{ m s}^{-1}$  spacing between  $10$  and  $20 \text{ m s}^{-1}$ , and  $2 \text{ m s}^{-1}$  spacing above  $20 \text{ m s}^{-1}$ . Note that all of the results presented in this manuscript are from observations taken between 25 August 2014 (when the most recent phase shift code for DPR was implemented) and 30 April

2015. Earlier observations used different phase shift codes and attenuator settings, which had some slight impact on the GMFs (not shown). The standard deviation in each bin is also calculated as is the correlation coefficient in the case of the matched KuPR-KaPR beams. The standard deviation serves as an implicit indicator of the quality of the derived GMF: Low values are desirable because they indicate that the 10m wind speed retrieved by GMI is sufficient to represent the sea state for the purposes of reproducing  $\sigma_0$ , and, when used in the combined framework, provide a stronger constraint on the PIA contributed by the precipitation column. The theoretical minimum standard deviation of  $\sigma_0$  for DPR, assuming the signal-to-noise ratio is large (true under almost all non-rain conditions), depends on the number of independent samples,  $N$ , taken. If the surface is modeled as a Rayleigh target (an incoherent sum from many specular points on the surface without any dominant scattering contribution) and a logarithmic receiver is used, then the standard deviation in dB is given by (Sauvageot 1992):

$$std(\sigma_0) = \frac{5.57}{\sqrt{N}}, \quad (3)$$

where  $N$  depends on incidence angle and varies between 100 and about 110 . Using these numbers, the nominal standard deviation in  $\sigma_0$ , from sampling alone, is a bit more than 0.5 dB.<sup>2</sup> Values higher than 0.5 dB could be caused by random errors in the GMI wind reference (this is compounded when the sensitivity of  $\sigma_0$  to wind is high) or that something other than wind speed is contributing the variation of  $\sigma_0$ , resulting in diminished impact of the  $\sigma_0$  observation on the precipitation retrieval. In Figure 3, the standard deviation of  $\sigma_0$  for the KuPR, in normal scan (NS) mode, and KaPR, in matched-scan (MS) and high-sensitivity (HS) modes, is shown as a function of DPR incidence angle for three wind speed bins centered on  $0.5 \text{ m s}^{-1}$ ,  $5 \text{ m s}^{-1}$ , and  $15 \text{ m s}^{-1}$ .

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<sup>2</sup>For off-nadir incidence, where there are multiple samples from the surface, a case can be made for integrating over all the data from the surface.

This should reduce the standard deviation of the  $\sigma_0$ ; however, in the DPR processing, the  $\sigma_0$  is based on the peak return power, not the integrated power.

242 At the low wind speed, the standard deviation is quite high (nearly 10 dB), particularly off nadir,  
243 but smaller (still 2-4 dB) near nadir at both frequencies (the lower Ku values are likely due to the  
244 saturation of the KuPR receiver). As the wind becomes calm, the surface is nearly specular and the  
245 sensitivity to small changes in wind speed is quite high off nadir, so random error in the reference  
246 wind is thought to primarily contribute to the large standard deviation there. Long-period swell  
247 also provides an increasing contribution to variation in  $\sigma_0$  (Tran et al. 2007) that is unrelated to  
248 the local wind speed. Finally, since the change in  $\sigma_0$  with respect to incidence angle is also high at  
249 low wind speeds, small changes in the incidence angle (the standard deviation of DPR incidence  
250 angle was around  $0.01^\circ$  in each angle bin) may also contribute to the high standard deviation at  
251 off-nadir angles.

252 At moderate and high wind speeds, the standard deviations are much lower and the pattern  
253 is shifted slightly to relatively high values near nadir and at the largest off-nadir angles, with  
254 minima around  $9^\circ$  for KuPR. Specular effects can again explain the near-nadir maximum, whereas  
255 the off-nadir maxima are likely a result of wind direction sensitivity (Wentz et al. 1984). The  
256 KaPR standard deviations are slightly higher for the MS than the HS data due to the shorter pulse  
257 width, and are qualitatively similar to the KuPR data. The effect of more stringent quality control  
258 (reduction of the cloud LWP, its spatial variability, and cost function thresholds by 50%; denoted  
259 QC2 in Figure 3) is also most evident here in reducing the KaPR standard deviation, but the  
260 differences are negligible enough (0.01 dB) that the original thresholds (QC1) are used to generate  
261 databases for the combined algorithm as this choice of thresholds provides more data, especially  
262 at higher wind speeds.

263 The two-dimensional GMFs of  $\sigma_0$  are shown in Figure 4. Most of the variability is exhibited  
 264 at low wind speeds at both Ku and Ka bands<sup>3</sup>. However,  $\sigma_0$  continues to decrease near nadir for  
 265 wind speeds as high as  $30 \text{ m s}^{-1}$ , which is approximately the upper limit of the reliable data that  
 266 has been collected so far. Off-nadir,  $\sigma_0$  appears to reach maxima at increasing wind speeds with  
 267 incidence angle. The standard deviation of  $\sigma_0$  reaches minima near the 0.5 dB sampling limit  
 268 at  $5\text{-}15^\circ$  and wind speeds between 5 and  $10 \text{ m s}^{-1}$ . There is also a minimum in the standard  
 269 deviation at Ku band (but not Ka band) at very low wind speeds near nadir. This is an artifact of  
 270 the saturation of the Ku receiver when  $\sigma_0 \geq 22.5 \text{ dB}$  (The Ka receiver saturates closer to 40 dB,  
 271 which is only observed over some land and ice surfaces). The higher standard deviations at the  
 272 off-nadir angles are likely a result of wind-direction induced variability in  $\sigma_0$ . In Figure 4f the  
 273 observed Ku-band  $\sigma_0$  is compared to the cutoff-invariant two-scale model (Soriano and Guérin  
 274 2008) using the Durden-Vesecky single-amplitude wave spectra (Durden and Vesecky 1985). This  
 275 model appears to produce a flatter  $\sigma_0$  when viewed with respect to incidence angle at low wind  
 276 speeds, but at winds above about  $8 \text{ m s}^{-1}$  has a comparable shape to the observed GMF minus a  
 277 small (1 dB) offset. These results are consistent with the comparisons of this model to airborne  
 278 observations of  $\sigma_0$  reported by Majurec et al. (2014).

279 The Ku-Ka  $\sigma_0$  correlation (Figure 4e) is an important component of the dual-frequency sur-  
 280 face reference technique (DSRT; Meneghini et al. (2012)). In the DSRT,  $\sigma_0$  is replaced by the  
 281 differential  $\sigma_0$ :

$$\delta\sigma_0 = \sigma_0(Ka) - \sigma_0(Ku) \quad (4)$$

282 and the method provides an estimate of the differential PIA,  $A(Ka)\text{-}A(Ku)$ . The errors in both  
 283 single-frequency SRT and DSRT methods are dominated by the fluctuations in the rain-free ref-

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<sup>3</sup>the Ka HS GMF is not shown, but is essentially identical to the MS data with a -0.2 dB offset owing to the inability of the larger pulse width  
 to capture the surface peak as effectively, especially near nadir.

284 erence data:  $\sigma_0$  and  $\delta\sigma_0$ . As the correlation between  $\sigma_0(\text{Ku})$  and  $\sigma_0(\text{Ka})$  increases, the variance  
 285 in  $\delta\sigma_0$  decreases so that the DSRT provides a potentially more accurate estimate of the path at-  
 286 tenuations. The correlations, which are near 0.8 in most DPR angle bins when all wind speeds  
 287 are considered, reduce to 0.1-0.4 for most wind speeds  $> 5 \text{ m s}^{-1}$  and off-nadir incidence angles.  
 288 This suggests that wind is responsible for most of the covariance in Ku and Ka  $\sigma_0$  but near-nadir  
 289 and at low winds the stronger correlations make the DSRT technique particularly useful.

### 290 **3. Combined Radar-Radiometer Retrieval of Precipitation and Surface Wind**

291 The MW emissivity model (optimized for GMI) and DPR wind- $\sigma_0$  GMFs described in section  
 292 2 are implemented in the forward modeling component of the GPM Combined Radar Radiometer  
 293 (CORRA) retrieval algorithm. A description of the radar component of this algorithm is given by  
 294 Grecu et al. (2011) and a more complete description of the algorithm architecture can be found in  
 295 the Algorithm Theoretical Basis Document (Olson and Masunaga 2015); for the purposes of this  
 296 manuscript, a brief summary and example case are presented in this section followed by validation  
 297 statistics. It is difficult to directly ascertain the improvement (if any) in rainfall estimates over  
 298 ocean owing to the lack of reliable direct measurements, but the algorithm can be assessed as to  
 299 how well the forward model matches GPM observations and buoy observations of wind speed.  
 300 The impact on retrieved precipitation amounts is also shown in this section.

#### 301 *a. Algorithm Description*

302 The CORRA algorithm uses an ensemble filter technique (Evensen 2006) to retrieve a set of pre-  
 303 cipitation profiles that are consistent with observations from KuPR, GMI, and KaPR (where avail-  
 304 able). The first step in this process is the creation of an ensemble of solutions that fit the observed  
 305 KuPR reflectivity profile without any consideration of the GMI, KaPR, or KuPR  $\sigma_0$  observations.

306 The randomly perturbed properties of each profile solution include the vertical profile of the hy-  
 307 drometer particle size distribution (PSD) intercept parameter ( $N_w$ ), degree of non-uniform beam  
 308 filling, the cloud liquid water profile, relative humidity, and 10m wind speed. For each solution,  
 309 the associated Ku and Ka  $\sigma_0$ , Ka reflectivities, and GMI radiances are calculated. The calculation  
 310 of Ka reflectivity's accounts for multiple scattering enhancements using the multiscatter library  
 311 developed by Hogan and Battaglia (2008).

312 The ensemble is then filtered using the observed Ku  $\sigma_0$ , GMI radiances, and Ka reflectivities and  
 313  $\sigma_0$  (where available). This is done by constructing an  $n_{var} \times n_{memb}$  vector  $\mathbf{X}_{ens}$  representing the  
 314 ensemble variables to be updated, including the perturbed variables, e.g.,  $N_w$  and 10m wind, and  
 315 derived/forward modeled variables, e.g., precipitation rate and brightness temperature. A separate  
 316  $n_{obs} \times n_{memb}$  vector  $\mathbf{Y}_{ens}$  consists of the forward modeled variables corresponding to the  $n_{obs} \times 1$   
 317 observation vector  $\mathbf{Y}_{obs}$  ( $\mathbf{R}$  is the corresponding observation error), which contains the observed  
 318  $\sigma_0$ , brightness temperatures, and Ka reflectivities. The ensemble state vector  $\mathbf{X}_{ens}$  is then updated  
 319 using the sample covariance:

$$\mathbf{X}_{ens} = \mathbf{X}_{ens} + \mathbf{Cov}_{XY}(\mathbf{Cov}_{YY} + \mathbf{R})^{-1}(\mathbf{Y}_{obs} - \mathbf{Y}_{ens}). \quad (5)$$

320 The algorithm output is derived from the updated ensemble and includes both mean and standard  
 321 deviations of the geophysical parameters of the ensemble and forward modeled observations. This  
 322 update is done separately for the Ku-only full swath (denoted as NS in GPM products) and Ku+Ka  
 323 inner swath (MS products).

### 324 *b. Example Case*

325 To illustrate the update process described by Eq. 5, the retrieval algorithm is applied to a GPM  
 326 overpass of a developing cyclone off the eastern coast of the United State on 26 January 2015 (Fig-



ure 5). This case provides an opportunity to examine the algorithm under a variety of precipitation and surface wind conditions.

The correlations (calculated from the initial, unfiltered ensemble) between the each observation type and the surface rain rate, as well as the correlations between each observation type and the 10m wind speed, are shown in Figure 6 for both radar frequencies and the horizontally-polarized GMI channels from 10-36 GHz (which are most sensitive to rain and wind over water surfaces). It is evident from these sensitivities that algorithm adjustments to precipitation rate in convective rain (echoes greater than 40 dBZ; purple colors in Figure 5) are mostly a response to the initial Ku and Ka  $\sigma_0$  error, whereas adjustments in stratiform rain are mostly a response to the Ka  $\sigma_0$  and GMI Tbs (note that in the heaviest rain, the correlation between rain rate and 36H Tb becomes negative as scattering dominates over emission). Note that in extremely heavy precipitation with large amounts of ice aloft, the variability of Ka  $\sigma_0$  due to multiple scattering begins to overwhelm the attenuation, and the correlation decreases. In these cases, the algorithm relies mostly on Ku  $\sigma_0$  to adjust the initial ensemble rain rates. In light and moderate rainfall, the 10m wind adjustment is mostly a response to Ku  $\sigma_0$ , especially away from the approximately  $9^\circ$  incidence angle at which Ku  $\sigma_0$  is insensitive to wind. Nevertheless there is some sensitivity of the 10 and 19 GHz radiances and Ka  $\sigma_0$  to wind under lighter precipitation. Due to the finite number of ensemble members, there are some spurious negative correlations between wind and the Tbs in heavier rain, but these are weak and do not substantially impact the output. The degree to which the ensemble spread is reduced after the filtering step is indicative of the overall information content in the observations for each variable of interest, and is provided as part of the standard CORRA output.

### 348 *c. Internal Validation*

349 Output from 400 GPM orbits between September 2014 and January 2015 are analyzed to assess  
350 the internal consistency between the forward model and observations before and after filtering.  
351 The mean bias and root-mean-square (rms) error between the initial ensemble mean and filtered  
352 ensemble mean for both NS (Ku+GMI) and MS (Ku+Ka+GMI) are given in Table 2. There is a  
353 general cold bias to the initial simulated brightness temperatures (Tbs) at all frequencies (although  
354 a warm bias is present in the 18 and 36 GHz channels at rain rates exceeding 10 mm hr<sup>-1</sup>). Both  
355 the rms error and magnitude of the bias are reduced after filtering as expected. The MS error and  
356 bias are larger than the NS error and bias because the initial ensemble profiles are constrained by  
357 the additional Ka band information and are less free to be adjusted to match the GMI radiances.  
358 In other words, the NS retrievals are over-fit to the Tbs, which suggests an increase in their error  
359 values in **R** is warranted.

360 The initial and filtered rms error and bias of  $\sigma_0$  is shown as a function of scan angle in Figure 7.  
361 There is a significant reduction in Ku rms error at all scan angles. The Ka error values are higher  
362 due to the stronger attenuation and multiple scattering effects, but errors are still reduced by nearly  
363 50% after the filtering step. The bias plots show a pattern of initial errors that are consistent with  
364 a low bias in the ENV wind (too high near nadir and too low off nadir). This bias appears to be  
365 more significant than any systematic bias in the precipitation attenuation, which would have the  
366 same sign regardless of scan angle.

### 367 *d. External Validation*

368 During September 2014-January 2015, 606 buoy observations from the ICOADS database were  
369 identified as being within 30 minutes of a GPM overpass and in the KuPR swath (308 of these

370 were within the KaPR swath) at the same time that DPR detected precipitation in the pixel nearest  
371 to the buoy location. These observations were used to validate the CORRA wind retrieval.

372 The wind rmse and bias are shown in Figure 8. Similar to the MERRA data analyzed in section  
373 2, these background winds are biased high below  $3 \text{ m s}^{-1}$  and biased low at higher wind speeds  
374 relative to the buoy observations. Root-mean-square errors increase from  $2 \text{ m s}^{-1}$  to  $4 \text{ m s}^{-1}$   
375 and NS errors are slightly higher than the MS or ENV errors. However, the bias is significantly  
376 reduced in the filtered datasets relative to the initial winds, indicating that while the retrievals are  
377 noisy, adjustments tend to be in the correct direction (this is consistent with the initial and filtered  
378  $T_b$  and  $\sigma_0$  biases as well).

379 The wind error is shown as a function of incidence angle in Figure 9. It is evident that the  
380 largest errors occur near the  $9^\circ$  incidence angle where there is little sensitivity of  $\sigma_0$  to wind  
381 speed (Figures 4 and 6 illustrate this behavior). Near nadir and beyond  $12^\circ$  incidence angles, the  
382 sensitivity is stronger and the wind errors are much smaller. The NS errors are similar and the  
383 MS errors are smaller than the  $4.26 \text{ m/s}$  error of ASCAT under raining conditions (Portabella et al.  
384 2012) and  $3.5 \text{ m/s}$  error Quikscat retrievals using a neural network to compensate for rain effects  
385 (Stiles and Dunbar 2010). These are also within the range of 2 to  $5 \text{ m/s}$  accuracy (depending on  
386 rain rate) of a globally-applicable rainy-atmosphere WindSat wind retrieval algorithm (Meissner  
387 and Wentz 2009). When stratified by rainfall rate, wind speed errors are similar for light ( $< 1$   
388  $\text{mm hr}^{-1}$ ) and moderate ( $1 \text{ mm hr}^{-1} < R < 10 \text{ mm hr}^{-1}$ ) precipitation rates, but increase at  
389 heavier precipitation rates as the wind-induced variability in  $\sigma_0$  and brightness temperatures is  
390 overwhelmed by the precipitation effects.

391 *e. Impact on Precipitation Retrieval*

392 Although the retrieval of wind in precipitation is useful for many applications, one of the main  
393 purposes of this work is to improve the precipitation retrieval by enforcing an internal consis-  
394 tency between the surface emissivity (which depends on wind) and observed  $\sigma_0$  which depends  
395 on both wind and precipitation-induced path-integrated attenuation (PIA). In this section we show  
396 the impact of switching from the SRT PIA (which infers PIA by comparing the observed  $\sigma_0$  to a  
397 reference outside the precipitation) to the coupled  $\sigma_0$ -emissivity model.

398 Theoretically, the use of  $\sigma_0$  as an observation (instead of SRT-derived PIA) should impact the  
399 agreement between observed and modeled Tbs in two ways: First, through adjustments to the rain  
400 column to match the observed  $\sigma_0$  by changing the PIA, and second, via changes in the surface  
401 emissivity. The relative importance of these mechanisms depends on the relative sensitivity of  
402 the Tbs and  $\sigma_0$  to changes in the rain column and surface wind. Figure 10 shows the change  
403 in near-surface precipitation rate retrieved by the GPM combined algorithm over ocean surfaces  
404 equatorward of  $55^\circ$  latitude (to eliminate possible sea ice) when the SRT PIA (single frequency for  
405 NS retrievals in top panels; DSRT in the MS retrieval shown in the bottom panels) is replaced with  
406 the observed  $\sigma_0$  in the observation vector. Light precipitation ( $< 1 \text{ mm hr}^{-1}$ ) is increased slightly  
407 in the NS swath, predominantly at wind speeds  $> 10 \text{ m s}^{-1}$  and at incidence angles less than  $12^\circ$ .  
408 The discontinuities in the  $10\text{-}12^\circ$  range are an artifact of the unavailability of the low-frequency  
409 GMI channels near the edge of the DPR swath (the deconvolution procedure requires coverage  
410 of the full footprint within the DPR swath). This suggests that GMI Tbs are driving the increase  
411 in precipitation, which is consistent with the weak Ku  $\sigma_0$ -precipitation correlation in light rain  
412 (Figure 6). Near the edges of the DPR swath, where the GMI Tbs are not used, there is not enough

413 information to significantly adjust the precipitation rate because the Ku-band PIA is small relative  
414 to the uncertainty in  $\sigma_0$ , so the SRT and coupled method have the same information content.

415 At moderate ( $1 \text{ mm hr}^{-1} < R < 10 \text{ mm hr}^{-1}$ ) precipitation rates, the wind- $\sigma_0$  correlation is  
416 still larger than the rain correlation at Ku band whereas Tbs are more sensitive to the precipitation  
417 (although there is still some wind sensitivity especially at 10H). This results in some compensating  
418 behavior, where it is “easier” for the algorithm to increase the wind speed to satisfy the Ku  $\sigma_0$   
419 observation but must reduce the precipitation rate to be consistent with the Tbs. In heavy rain ( $>$   
420  $10 \text{ mm hr}^{-1}$ ), the ensemble variance in  $\sigma_0$  and the Tbs is dominated by variance in the rain column,  
421 rather than surface wind, and where both observations are available only a very small reduction  
422 in precipitation is noted with the coupled forward model relative to the SRT method. When only  
423 Ku  $\sigma_0$  is available in the outer swath, however, there is a reduction in precipitation relative to the  
424 SRT version. The mean precipitation rate from the coupled model is more consistent across the  
425 different scan angles than the SRT version (not shown) which suggests that the SRT PIA may be  
426 biased high at the off-nadir angles and wind speeds from  $5\text{-}10 \text{ m s}^{-1}$ .

427 The Ku-Ka (MS) retrievals are generally more stable when comparing the SRT and coupled  
428 versions of the algorithm, but some changes are still notable. The increase in light precipitation  
429 is still present, but moderate and heavy precipitation show some different behavior from the NS  
430 retrievals with increases in light winds (below about  $5 \text{ m s}^{-1}$ ) and little change at higher wind  
431 speeds. There is not much sensitivity of Tb to wind at low wind speeds, so this appears to be  
432 driven by an increase in the inferred PIA in the coupled model relative to the dual-frequency SRT.

#### 433 **4. Summary**

434 The Global Precipitation Measurement core satellite launched in February, 2014 carries a passive  
435 microwave imager (GMI) and dual-frequency radar (DPR) designed specifically to provide the

most accurate instantaneous precipitation estimates currently available from space and serve as a reference for precipitation retrievals from other passive microwave imagers with similar channel sets (Kummerow et al. 2015). The GPM combined algorithm plays a key role in this process by providing precipitation estimates that are consistent with both GMI and DPR measurements. This algorithm uses physically-based forward models to simulate GMI and DPR measurements and it is desirable that those models use the same geophysical input parameters wherever possible.

This study explored the feasibility of using internally consistent relationships between wind, emissivity, and backscatter for water surfaces in the combined algorithm. We first evaluated the FASTEM 4/5 (Liu et al. 2011) and Meissner and Wentz (2012) emissivity models in a GMI-only non-precipitation retrieval against buoy observations obtained from the ICOADS dataset. The Meissner-Wentz model provided the lowest root-mean-square error ( $1.3 \text{ m s}^{-1}$ ) and was used to create a geophysical model function (GMF) for DPR Ku and Ka  $\sigma_0$  as a function of 10m wind speed and incidence angle by matching the GMI retrievals to DPR observations under clear conditions.

The Meissner-Wentz emissivity model and DPR GMFs were then implemented in the GPM combined algorithm. This coupled forward model indicated that the sensitivity of  $\sigma_0$  to wind at Ku band dominates the precipitation sensitivity particularly in light to moderate rain and at low wind speeds, where the brightness temperatures are more sensitive to precipitation (although there is still some wind sensitivity, particularly at 10 and 18 GHz at horizontal polarization in light and shallow precipitation). Therefore, the surface reference (SRT) estimate of the DPR path-integrated attenuation (PIA) was replaced with  $\sigma_0$  in the observation vector. This is desirable because  $\sigma_0$  is directly observed by DPR while the SRT PIA includes implicit assumptions and can be unphysically negative in light rain. Because  $\sigma_0$  depends on both the 10m wind speed and

attenuation from atmospheric gases, clouds, and precipitation, the 10m wind speed was added to the retrieval state vector.

The combined wind/precipitation retrievals were then evaluated against the ICOADS buoy dataset under precipitating conditions, which have been a challenge for surface wind retrievals from standalone passive radiometers (e.g., WindSat) or scatterometers. Although the retrievals were noisier than under clear conditions (rmse of  $3.7 \text{ m s}^{-1}$  for Ku+GMI and  $3.2 \text{ m s}^{-1}$  for Ku+Ka+GMI), there was a significant reduction in the bias from the background data provided by GANAL (-10%) to the Ku+GMI (-3%) and Ku+Ka+GMI (-5%) retrievals. The impact on precipitation retrievals was also evaluated. Ku+GMI retrievals of precipitation increased slightly on the light end ( $< 1 \text{ mm hr}^{-1}$ ) and decreased in moderate to heavy precipitation ( $> 1 \text{ mm hr}^{-1}$ ) due to compensating effects of wind on  $\sigma_0$  and emissivity requiring changes in the precipitation column to maintain consistency with the observations. The Ku+Ka+GMI retrievals, being additionally constrained by the Ka reflectivity, did not change as much although a slight increase in moderate and heavy precipitation at low wind speeds was noted.

While GPM was not designed specifically to measure ocean surface winds, this study demonstrates that such measurements are quite feasible in clear-sky conditions. In precipitation, using a coupled emissivity-backscatter GMF produces reasonable results that achieve the goal of internal consistency in the combined algorithm. The results presented here should only be considered as a proof of concept, as additional details that we did not consider, such as wind direction, the effect of rain on the scattering properties of water surfaces, and spatial correlation of the wind field, are left to future work.

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661 and 10m wind and MERRA atmospheric parameters. No offsets were applied to the 183 GHz channels.

	FASTEM4		FASTEM5		Meissner-Wentz	
Channel	bias	rmse	bias	rmse	bias	rmse
10.65V	1.6	0.8	0.6	0.7	-1.0	0.7
10.65H	2.9	0.9	-0.4	0.9	-0.7	0.9
18.7V	0.6	1.0	0.2	1.0	-0.5	1.0
18.7H	2.3	1.6	0.3	1.5	0.1	1.5
23.8V	-0.5	1.5	-0.5	1.4	-0.6	1.4
36.64V	0.6	0.9	0.6	0.9	0.6	0.9
36.64H	2.9	1.6	2.1	1.6	1.3	1.5
89V	-0.4	1.0	0.0	1.0	0.8	1.1
89H	1.5	2.2	2.5	2.2	1.7	2.2
166V	-0.1	1.4	-0.3	1.4	-0.3	1.4
166H	0.1	2.8	0.2	3.1	0.3	3.2
183 $\pm$ 3	-2.8	3.7	-2.8	3.6	-3.0	3.6
183 $\pm$ 7	-0.7	1.8	-0.9	1.9	-1.0	1.9

662      TABLE 2. Root-mean-square error and bias of ensemble-mean deconvolved GPM Microwave Imager (GMI)  
663      radiances before and after filtering.

	10V	10H	18V	18H	23V	36V	36H	89V	89H
Initial rmse (K)	6.7	11.3	11.6	19.8	12.0	15.5	26.6	23.1	25.9
Initial Bias (K)	-3.7	-7.2	-6.0	-10.9	-6.5	-9.0	-14.3	-16.1	-16.0
NS rmse (K)	4.6	7.1	7.2	11.1	8.1	9.9	15.1	16.2	16.9
NS Bias (K)	-1.7	-3.6	-2.9	-5.0	-3.3	-4.8	-5.7	-11.6	-8.2
MS rmse (K)	8.1	9.8	11.1	15.1	12.7	13.6	20.0	22.0	24.8
MS Bias (K)	-2.5	-4.0	-4.4	-6.8	-5.2	-5.9	-8.4	-13.2	-12.5

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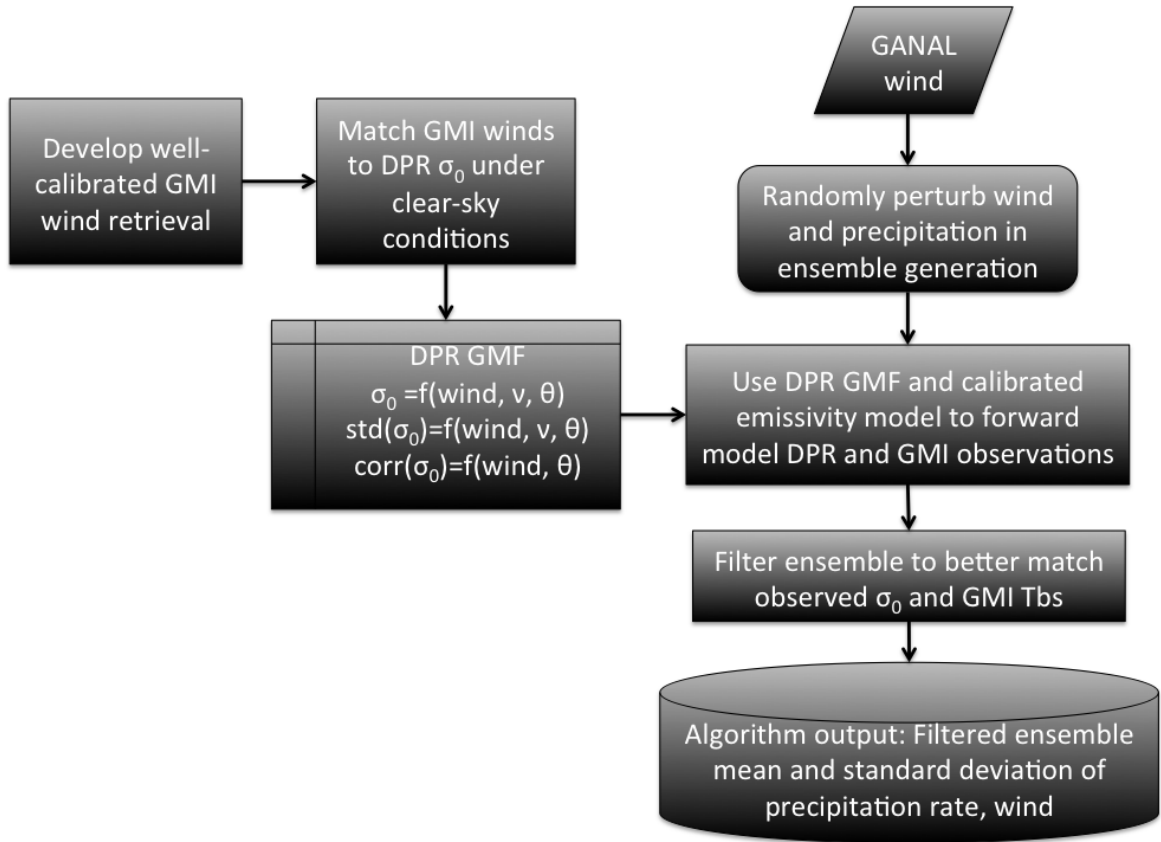


FIG. 1. Flow chart of the process by which the DPR geophysical model function (GMF) is derived and used by the combined DPR-GMI precipitation algorithm.

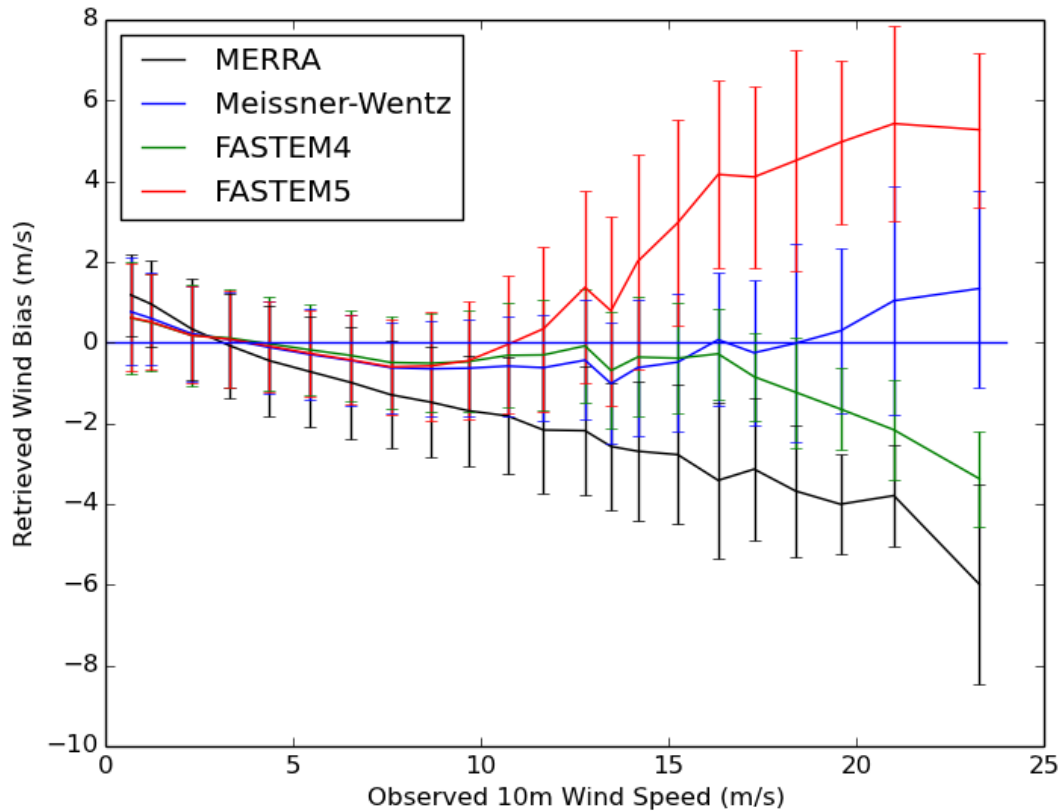


FIG. 2. MERRA and GMI-retrieved wind speed bias relative to ICOADS buoy observations from March-December 2014. Error bars represent 1 standard deviation of the difference between observed and retrieved wind speeds in each bin.

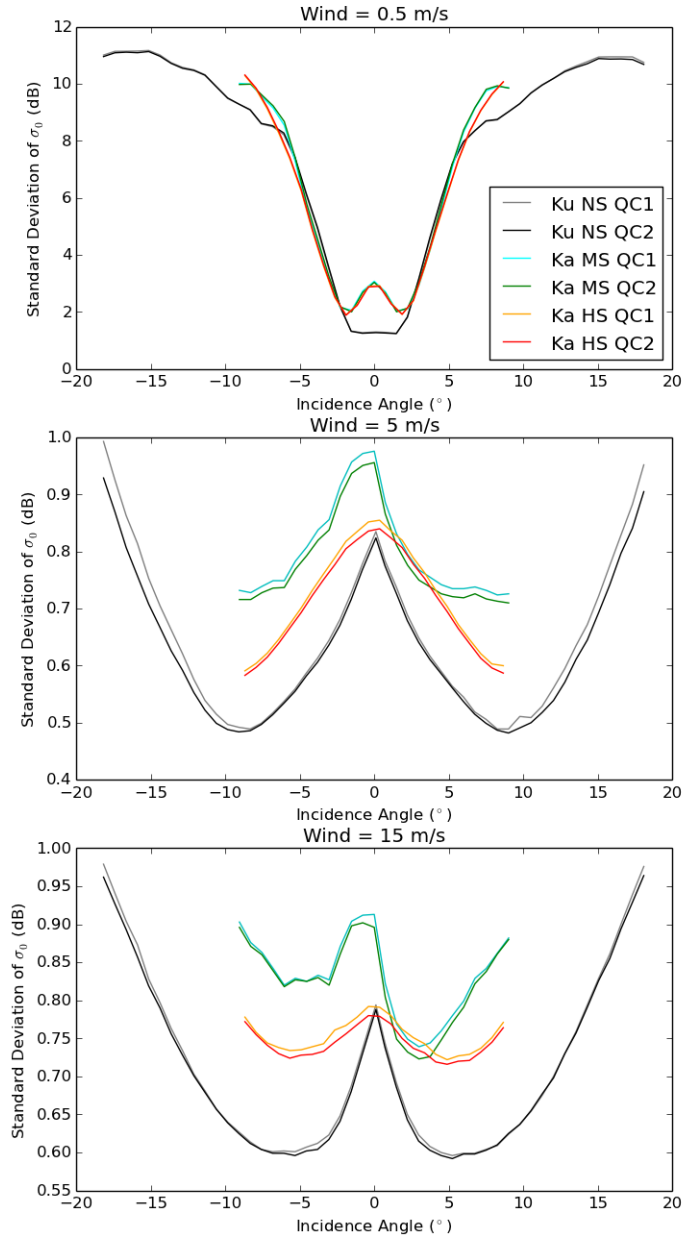


FIG. 3. Standard deviation of  $\sigma_0$  in three wind speed bins:  $0.5 \text{ m s}^{-1}$  (top),  $5 \text{ m s}^{-1}$  (middle), and  $15 \text{ m s}^{-1}$  (bottom). The different colors represent different frequencies, DPR modes (NS = Normal Scan, MS = Matched Scan, HS = High Sensitivity), and quality control of the reference wind.

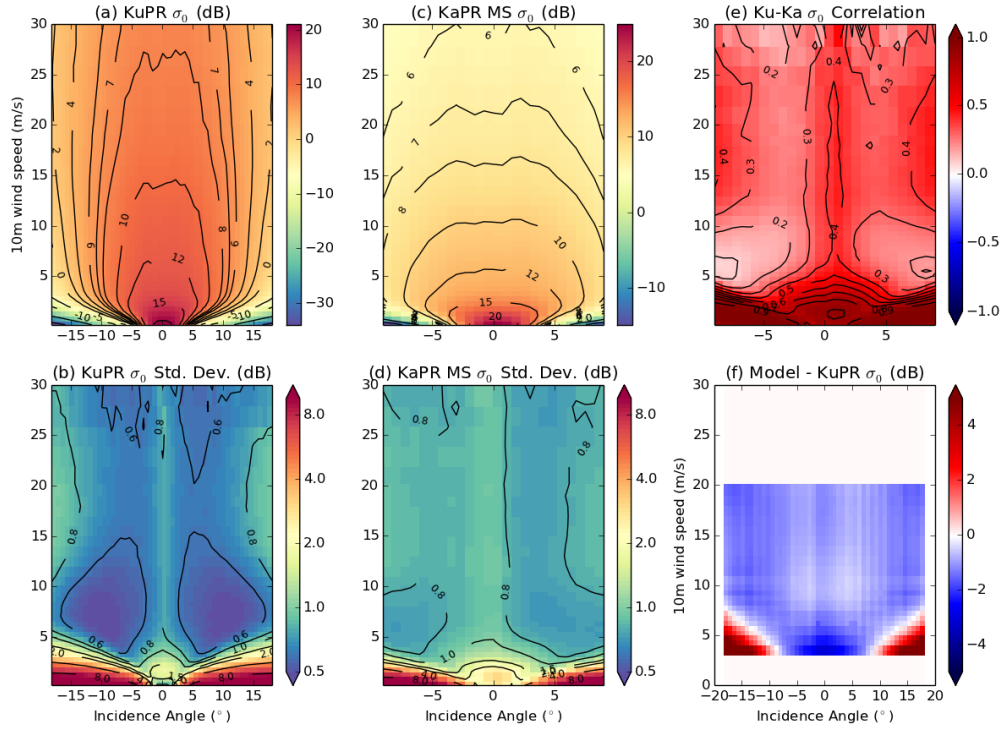


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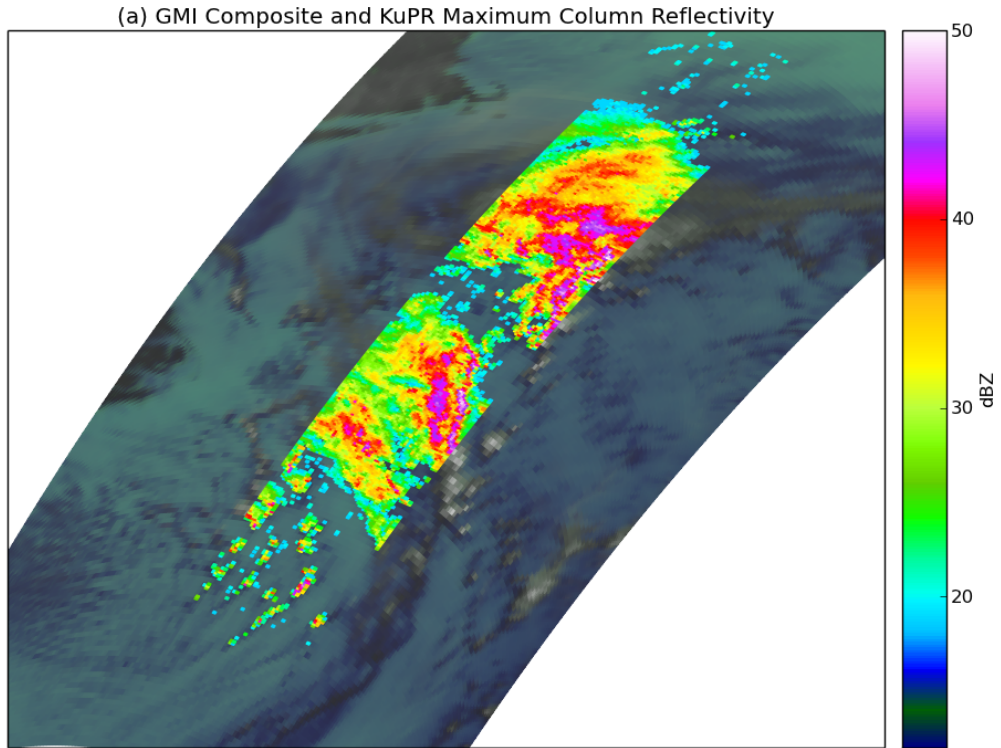


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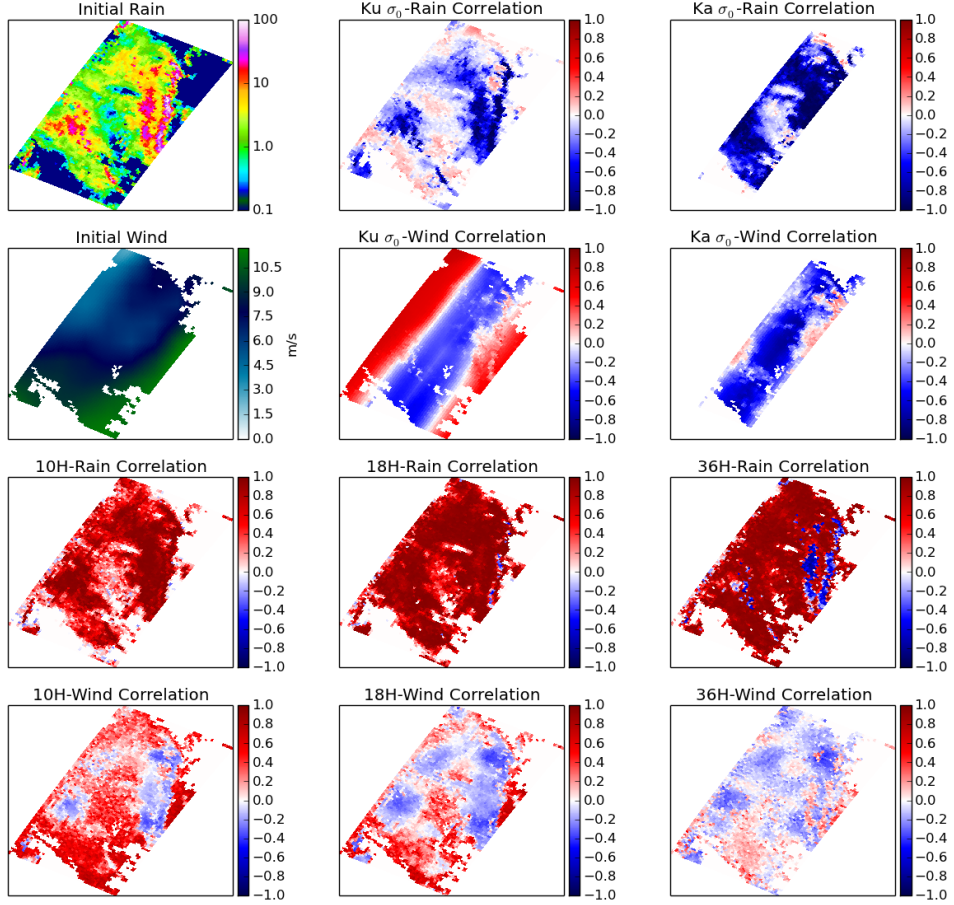


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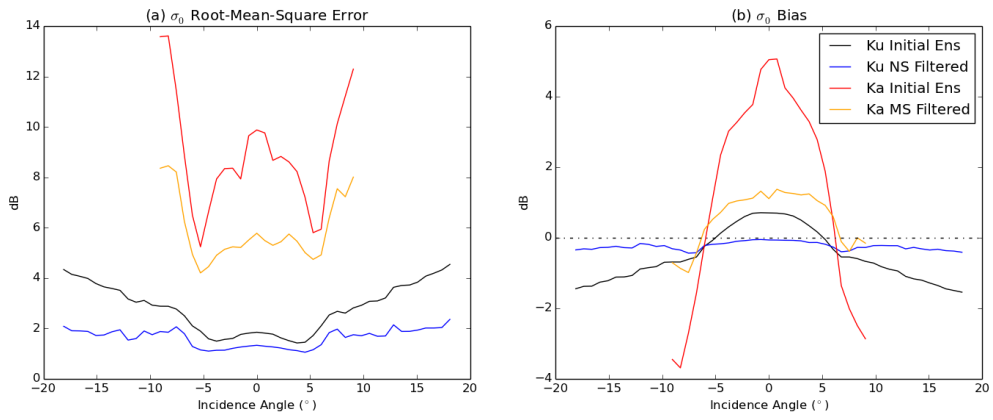


FIG. 7. Root-mean-square error (a) and bias (b) of the initial and filtered ensemble mean  $\sigma_0$  as a function of incidence angle.

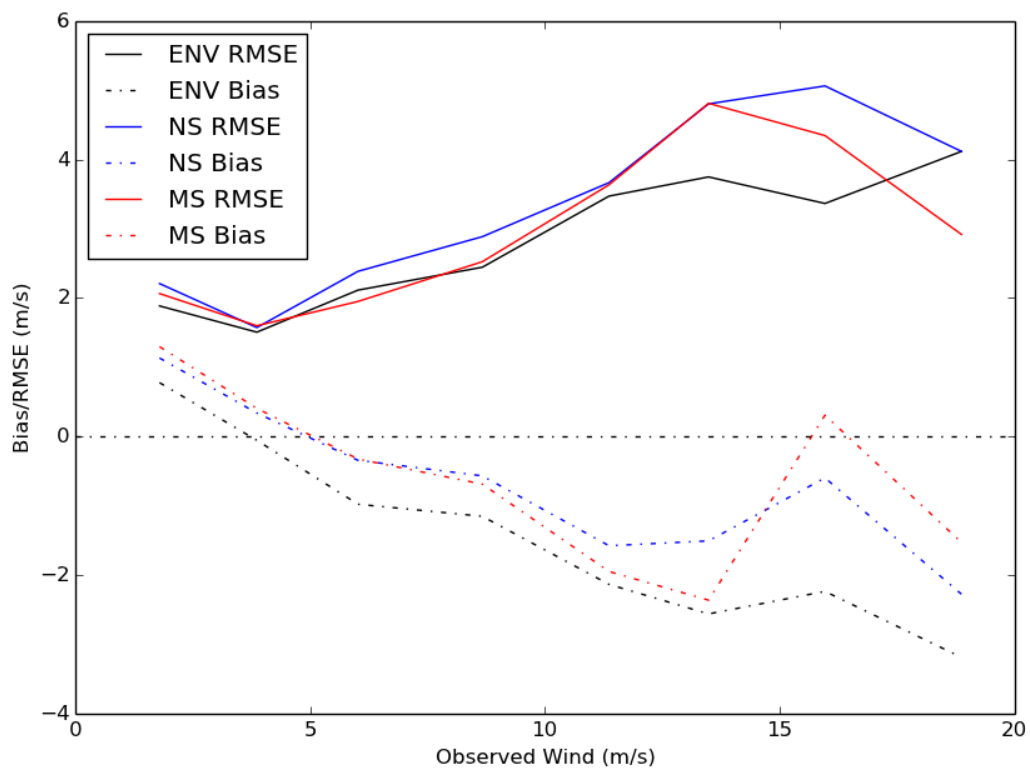


FIG. 8. Background (JMA Global Analysis GANAL; 2A-ENV) and retrieved wind root-mean-square error and bias relative to ICOADS buoy observations in precipitating pixels.

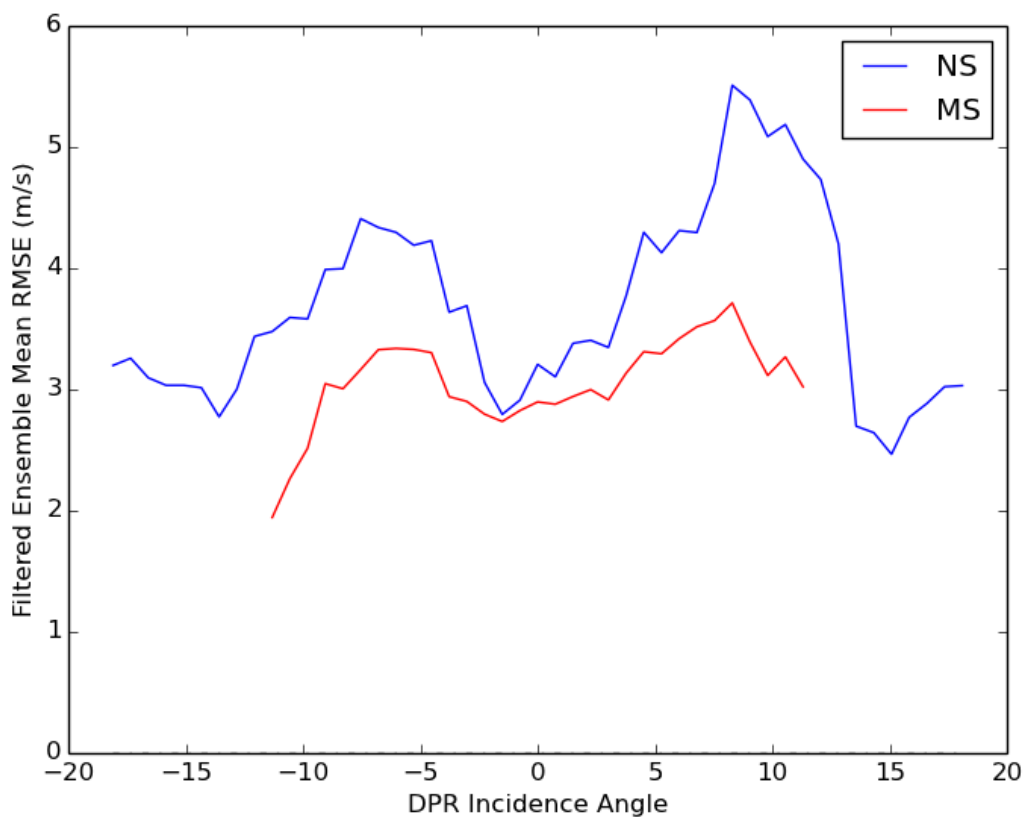


FIG. 9. Retrieved wind root mean square error as a function of DPR incidence angle. The data are smoothed using a 7-bin centered average in order to reduce noise from the small sample size in each angle bin.

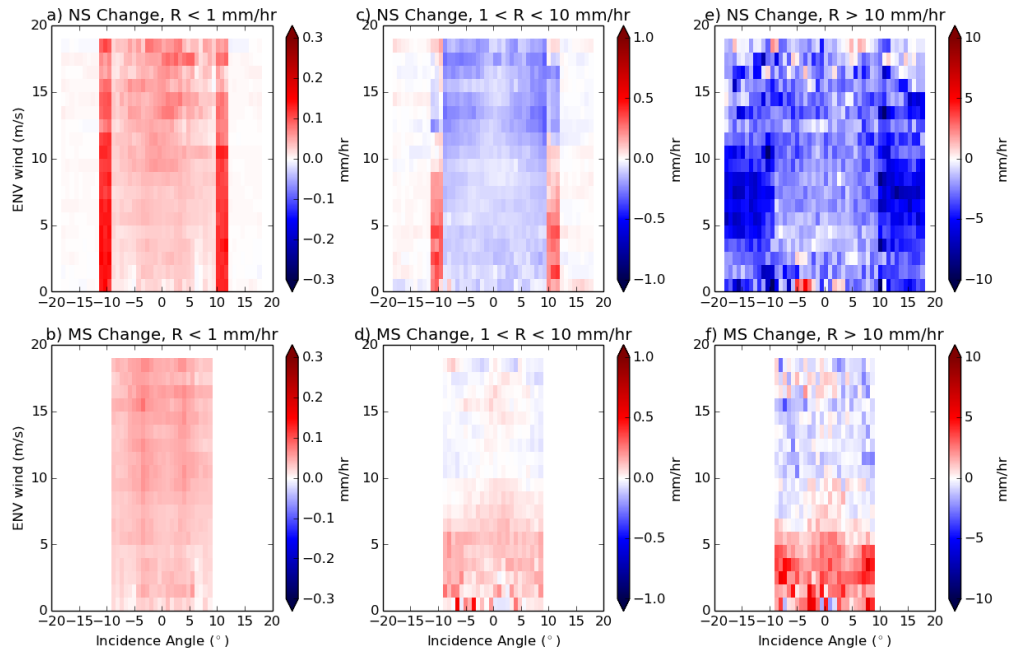


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